ON SUPERSYMMETRIC REGULARIZATION in FIELD THEORY and HOLOGRAPHY

Davide Cassani LPTHE — Université Paris 6

Supersymmetric theories, dualities and deformations

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Based on: 1606.02724 with Benetti Genolini, Martelli, Sparks + to appear + previous work with Assel, Di Pietro, Komargodski, Lorenzen, Martelli

Supersymmetric regularization

- Recently: many new exact results in supersymmetric QFT mainly thanks to localization
- need to renormalize UV divergences



- → crucial to use a regularization scheme that preserves supersymmetry
 - One way to assess this : check susy Ward identities at the end
- Even if regularization is supersymmetric, there may be ambiguities,
 - → important to classify them in order to extract physically meaningful result

Example

Gerchkovitz, Gomis, Komargodski

Partition function of N=2 SCFT's on S^4 as a function of marginal couplings τ , $\bar{\tau}$ computes the Kähler potential $K(\tau, \bar{\tau})$ on the conformal manifold. Ambiguous by $F(\tau) + \overline{F}(\bar{\tau})$: correspond to Kähler transformations.

Field theory side

• Consider a 4d, N=1 SCFT.

can be placed on a space with topology preserving some susy



 \leftarrow Hopf surfaces $\mathcal{H}_{p,q}$ with complex structure parameters p, q

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susy Ward identities : $\delta_{ ext{Hermitian}}(-\log Z) = 0$

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• Localization yields: $Z[\mathcal{H}_{p,q}] = e^{-\mathcal{F}(p,q)}\mathcal{I}(p,q)$ Assel, D.C., Martelli only depends on p,qalthough we allowed for a general metric $\mathcal{I}(p,q) = \operatorname{tr}\left[(-1)^F p^{J+J'-\frac{R}{2}} q^{J-J'-\frac{R}{2}}\right]$ susy Ward identities

supersymmetric Casimir energy

Assel, D.C., Martelli $Z[\mathcal{H}_{p,q}] = e^{-\mathcal{F}(p,q)} \mathcal{I}(p,q)$ Asset, D.C., N $\int uperconformal index$ $p = \mathrm{e}^{-eta b_1}, \quad q = \mathrm{e}^{-eta b_2}$ $\mathcal{F}(p,q) = \beta E_{susv}(b_1,b_2)$ $E_{\text{susy}} = \frac{2}{3} \left(b_1 + b_2 \right) \left(a - c \right) + \frac{2}{27} \frac{\left(b_1 + b_2 \right)^3}{b_1 b_2} \left(3 c - 2 a \right)$ ightarrow supersymmetric Casimir energy a,c central charges

$$Z = \int \mathcal{D}\phi \, e^{-S[\phi]} = \operatorname{Tr} e^{-\beta H}$$

$$S^{1}_{\beta}$$
 X

 $Z \sim e^{-eta E_{ ext{Casimir}}}$ as $eta
ightarrow \infty$

(projects to ground state)

The importance of being supersymmetric

 $Z_{non-susy}$ is ambiguous due to local counterterm :

$$-\log Z_{\text{non-susy}} + b \int d^4x \sqrt{g} R^2$$

$$\neq 0 \text{ on } S^1 \times S^3$$

→ Casimir energy on $S^1 \times S^3$ is scheme-dependent

In susy case :

ambiguities are gauge-invariant (new minimal) supergravity actions of dim = 4

- they are all F-terms
- F-terms vanish on susy backgrounds with 2 supercharges of opposite R-charge

$$-\log Z_{\text{susy}} + b \int d^4x \sqrt{g} \left[(R + 6V_{\mu}V^{\mu})^2 - 8F_{\mu\nu}F^{\mu\nu} \right] + \text{fermions}$$
$$= 0 \text{ on } S^1 \times S^3 !$$

The importance of being supersymmetric



• superalgebra implies the BPS condition $E_{susy} = \frac{1}{r} \langle R \rangle$ on round $S^1_\beta \times S^3_r$

→ vacuum charged!

gauge/gravity

QFT

gravity string theory

Credit: Next Architects

Holographic dictionary



Holographic Casimir energy

$$Z[\mathcal{H}_{p,q}] = e^{-\mathcal{F}(p,q)} \mathcal{I}(p,q)$$
superconformal index

 $\mathcal{F}(p,q) = \beta E_{susy}(b_1,b_2)$

$$E_{
m susy} = rac{2}{3} \left(b_1 + b_2
ight) \left(a - c
ight) + rac{2}{27} rac{(b_1 + b_2)^3}{b_1 b_2} (3 \, c - 2 \, a)$$

dominates Z at large N_c \rightarrow prediction for dual supergravity solutions

• There should be a family of susy solutions with $\,\partial M_5 = {\cal H}_{p,q}\,$ such that :

$$S_{
m 5d\,sugra}[M_5] = rac{2}{27}etarac{(b_1+b_2)^3}{b_1b_2}rac{\pi^2\ell^3}{\kappa_5^2}$$

$$c=a=\pi^2\ell^3/\kappa_5^2$$

Holographic Casimir energy

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superconformal index

 $\mathcal{F}(p,q) = \beta E_{susy}(b_1,b_2)$

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dominates Z at large N_c \rightarrow prediction for dual supergravity solutions

• There should be a family of susy solutions with $\,\partial M_5 = {\cal H}_{p,q}\,$ such that :

$$S_{5d \text{ sugra}}[M_5] = \frac{2}{27} \beta \frac{(b_1 + b_2)^3}{b_1 b_2} \frac{\pi^2 \ell^3}{\kappa_5^2} = \frac{16}{27} \frac{\beta}{r} \frac{\pi^2 \ell^3}{\kappa_5^2}$$
$$c = a = \pi^2 \ell^3 / \kappa_5^2 \qquad b_1 = b_2 = \frac{1}{r} \text{ round } S_\beta^1 \times S_r^3$$

On-shell action a priori divergent

$$S = rac{1}{2\kappa_5^2} \int_{M_5} d^5 x \sqrt{g} \left(R[g] - F^2 + rac{12}{\ell^2}
ight) - A \wedge F \wedge F$$

5d supergravity action S_{bulk}

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 $+ rac{1}{\kappa_5^2} \int_{\partial M_5} d^4x \sqrt{h} igg(K - rac{3}{\ell} - rac{\ell}{4} R[h]igg) \, .$

5d supergravity action S_{bulk}

divergent counterterms

On-shell action a priori divergent

$$\begin{split} S &= \frac{1}{2\kappa_5^2} \int_{M_5} d^5 x \sqrt{g} \left(R[g] - F^2 + \frac{12}{\ell^2} \right) - A \wedge F \wedge F & \text{5d supergravity} \\ &+ \frac{1}{\kappa_5^2} \int_{\partial M_5} d^4 x \sqrt{h} \left(K - \frac{3}{\ell} - \frac{\ell}{4} R[h] \right) & \text{divergent counterterms} \\ &+ \frac{1}{\kappa_5^2} \int_{\partial M_5} d^4 x \sqrt{h} \left(\lambda_1 R_{ijkl}[h]^2 + \lambda_2 R_{ij}[h]^2 + \lambda_3 R[h]^2 + \lambda_4 F_{ij}^2 \right) \end{split}$$

finite boundary terms -> parameterize different schemes

On-shell action a priori divergent

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finite boundary terms -> parameterize different schemes

• For round $S_{\beta}^{1} \times S_{r}^{3}$ the most natural candidate is pure Anti de Sitter space

Choosing $\lambda_1 = \lambda_2 = \lambda_3 = 0$

$$S=rac{3}{4}rac{eta}{r}rac{\pi^2\ell^3}{\kappa_5^2} \hspace{0.1in}
eq rac{16}{27}rac{eta}{r}rac{\pi^2\ell^3}{\kappa_5^2}$$

does **NOT match** the susy QFT result!

...could think about adjusting the λ 's ... but ...

• $\frac{1}{\sqrt{g^{bdy}}} \frac{\delta S}{\delta A_i^{bdy}} = j^i$ holographic R-current \rightarrow holographic R-charge $\langle R \rangle$

however unavoidably for AdS_5 : $\langle R \rangle = 0$

 \Rightarrow $\langle E \rangle \neq \frac{1}{r} \langle R \rangle$ BPS relation violated

unless one tunes λ 's such that

 $\langle E \rangle = \frac{1}{r} \langle R \rangle = 0$ misses Casimir energy

...something is not working here

New supergravity solutions

Benetti Genolini, D.C., Martelli, Sparks

We constructed a very general AIAdS solution perturbatively near the boundary

$$g^{
m bulk} = rac{{
m d}
ho^2}{
ho^2} + rac{1}{
ho^2} \left[g^{
m bdy} + g^{(2)}
ho^2 + \left(g^{(4)} + ilde{g}^{(4)} {
m log}
ho^2
ight)
ho^4 + ...
ight]$$



graviphoton field A^{bulk} also determined

 $A^{\mathrm{bdy}} = -\frac{1}{\sqrt{3}} \left[\frac{\mathrm{i}}{8} u \,\mathrm{d}\tau + \frac{1}{4} u (\mathrm{d}\psi + a) + \frac{\mathrm{i}}{4} (\partial_{\bar{z}} w \,\mathrm{d}\bar{z} - \partial_{z} w \,\mathrm{d}z) + (\gamma \mathrm{d}\psi \right]$

4 free subleading functions $k_1(z, \overline{z}), k_2(z, \overline{z}), k_3(z, \overline{z}), k_4(z, \overline{z})$

New supergravity solutions

non-trivial:

susy involves solving 6th-order equation for auxiliary Kähler metric

$$abla^2\left(rac{1}{2}
abla^2 R+rac{2}{3}R_{pq}R^{pq}-rac{1}{3}R^2
ight)+
abla^m(R_{mn}\partial^n R)\ =\ 0$$
D.C., Lorenzen, Martelli

bulk metric has complex components (but real in Lorentzian signature)





on-shell action gauge-dependent due to Chern-Simons term $A \wedge F \wedge F$

- \rightarrow crucial to fix the gauge properly
 - Killing spinors independent of time

gauge field globally well-defined

Vary the boundary data keeping complex structure on ∂M fixed

 \leftrightarrow vary the functions $u(z, \overline{z})$, $w(z, \overline{z})$ with globally def. variations

$$\delta S = \int_{M_4}\!\!\mathrm{d}^4x \sqrt{g^{\mathrm{bdy}}} \left(frac{1}{2} T^{ij} \delta g^{\mathrm{bdy}}_{ij} + j^i \delta A^{\mathrm{bdy}}_i
ight)$$

 Λ There is no choice of $\lambda_1, \lambda_2, \lambda_3, \lambda_4$ such that $\delta_u S = 0 = \delta_w S$

→ Holographic renormalization violates field theory susy Ward identities!

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new boundary term ΔS_{new} such that $S_{susy} = S + \Delta S_{new}$ satisfies $\delta S_{susy} = 0$

$$egin{aligned} \Delta S_{
m new}[M_4] &= rac{1}{\kappa_5^2} \int_{M_4} (\mathrm{i} A^{
m bdy} \wedge \Phi + \Psi) & S ext{ is taken with } \ \lambda_1 &= \lambda_2 = \lambda_3 = \lambda_4 = 0 \end{aligned}$$
 $\Phi &= rac{1}{72\sqrt{3}} \left(u^3 + 4u \Box w
ight) \mathrm{i} \, \mathrm{e}^w \mathrm{d} z \wedge \mathrm{d} ar{z} \wedge (2 \, \mathrm{d} \psi + \mathrm{i} \, \mathrm{d} au) \end{aligned}$
 $\Psi &= rac{1}{2^{11} 3^2} \left(-19 u^4 - 48 u^2 \Box w
ight) \mathrm{d}^4 x \sqrt{g^{
m bdy}} \end{aligned}$

We propose that $S_{susy}[M_5] = S[M_5] + \Delta S_{new}[M_4]$

should be identified with the SCFT $-\log Z_{susy}[M_4]$

Although we don't know the solution in the interior, under topological assumption we can actually evaluate the on-shell action



• S_{bulk} reduces to a boundary term

• 4 subleading functions $k_1(z, \overline{z}), k_2(z, \overline{z}), k_3(z, \overline{z}), k_4(z, \overline{z})$ drop out !

$$S_{
m susy} \;=\; S + \Delta S_{
m new} \;=\; rac{\gamma^2}{27} \int_{M_4} \!\!\!\mathrm{d}^4 x \, \sqrt{g^{
m bdy}} \, R_{(2d)} \, rac{\ell^3}{\kappa_5^2}$$

$$8\pi^2etarac{b_1+b_2}{b_1b_2}$$

 $S_{
m susy}\ =\ S+\Delta S_{
m new}\ =\ rac{\gamma^2}{27}\int_{M_4}\!\!\!\mathrm{d}^4x\,\sqrt{g^{
m bdy}}\,R_{(2d)}\,rac{\ell^3}{\kappa_5^2}$



from $A^{bdy} = \dots + \gamma d\psi$ $\gamma = \frac{1}{2}(b_1 + b_2)$ $8\pi^2 \beta \frac{b_1 + b_2}{b_1 b_2}$ $S_{susy} = S + \Delta S_{new} = \frac{\gamma^2}{27} \int_{M_4} d^4 x \sqrt{g^{bdy}} R_{(2d)} \frac{\ell^3}{\kappa_5^2}$ $= \frac{2}{27} \beta \frac{(b_1 + b_2)^3}{b_1 b_2} \frac{\pi^2 \ell^3}{\kappa_5^2}$



Conclusions

- Standard holographic renormalization in 5d violates susy
- Identified boundary terms ΔS_{new} that restore susy Ward identities
- Constructed asymptotic solutions such that

$$S_{\text{susy}} = S + \Delta S_{\text{new}} = \frac{2}{27} \beta \frac{(b_1 + b_2)^3}{b_1 b_2} \frac{\pi^2 \ell^3}{\kappa_5^2}$$
 iield theory

? △S_{new} in covariant form?
? which boundary fields involved?

First principle derivation such that bulk + boundary action is supersymmetric?

→ Zaffaroni's talk

• does ΔS_{new} have consequences for susy AdS5 black holes?

... thank you !